

# Beam Instability and Microbunching due to Coherent Synchrotron Radiation

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## Introduction

- A relativistic electron beam moving on a circular orbit in free space can radiate coherently at the wavelengths that exceed the length of the bunch.
- Coherent radiation at shorter wavelengths can result from density fluctuations in the beam with characteristic length much shorter than the bunch length.
- If the radiation reaction force drives the growth of the initial fluctuation, one can expect an instability which leads to micro-bunching of the beam and increased coherent radiation at short wavelengths.

## Mechanism of the instability

Let us assume a small initial sinusoidal density perturbation on the beam,  $\delta n = \epsilon \sin kz$

- Due to the CSR wake,  $\delta n$  induces energy modulation in the beam  $\delta E_1$
- Momentum compaction of the ring translates  $\delta E_1$  into  $\delta n$ . Under certain conditions, the final  $\delta n$  is greater than the initial one.
- Energy spread introduces Landau damping and stabilizes short wavelengths.
- Wall shielding of CSR and finite length of the bunch limits the instability at large wavelength.
- Transverse beam emittance mixes the particle over the wavelength and may have a stabilizing effect on the instability

## Theory (Heifets, Stupakov, 2002)

Since we are interested in the wavelength much shorter than the bunch length, consider a *coasting* beam moving in a circular orbit of radius  $R$  in free space.

$\rho(\delta, z, s)$  — longitudinal distribution function,  $dN = dz \int \rho(\delta, z, s) d\delta$ .

Vlasov equation (neglect damping and quantum diffusion)

$$\frac{\partial \rho}{\partial s} - \eta \delta \frac{\partial \rho}{\partial z} - \frac{r_0}{\gamma} \frac{\partial \rho}{\partial \delta} \int_{-\infty}^{\infty} dz' d\delta' W(z - z') \rho(\delta', z', s) = 0$$

$$\delta = \Delta E / E$$

$$s = ct$$

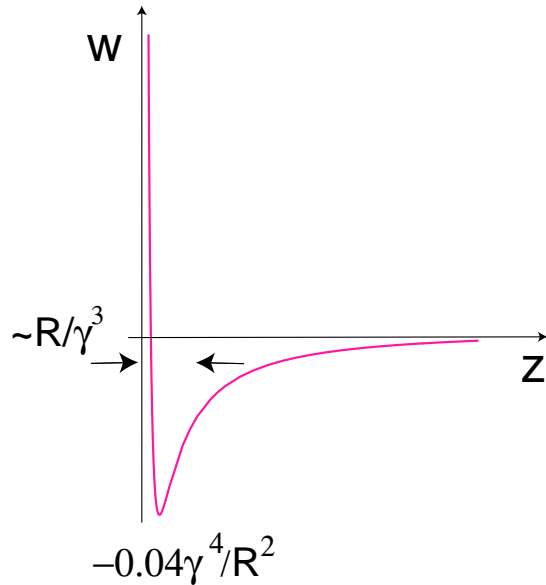
$\eta$  — slip factor

$r_0$  — classical electron radius

$W(z - z')$  — wake function (per unit length of the path).

## CSR wake

A relativistic particle moving in vacuum in a circular orbit of radius  $R$ , in steady state, generates a CSR wake (per unit length of path) (Murphy et al., 1995; Derbenev et al. 1995)



$$W(z) \approx -\frac{E_{\parallel}}{q} = -\frac{2}{3^{4/3} R^{1/3} z^{4/3}}$$

This wake neglects transient effects at the entrance to and exit from the magnet. It is valid if ( $\lambda$  – wavelength)

$$l_{\text{magnet}} > \lambda^{1/3} R^{2/3}$$

- For  $R \approx 30$  m and  $\gamma \approx 10^4$ ,  $R/\gamma^3 = 3 \cdot 10^{-11}$  m—negligibly small
- Shielding effects become important at the distance  $z \gtrsim \frac{R^{1/2}}{a^{3/2}}$ ,  $a$  – gap between walls

Neglect the shielding effect, and assume a steady-state wake

$$W(z) = \frac{2}{(3R^2)^{1/3}} \frac{\partial}{\partial z} \frac{1}{z^{1/3}} \quad \text{for } z > 0,$$

and  $W(z) = 0$  for  $z \leq 0$ . The radiation wakefield is localized in front of the moving charge.

Impedance

$$Z(k) = \frac{1}{c} \int_0^\infty dz W(z) e^{-ikz} = iA \frac{k^{1/3}}{cR^{2/3}}.$$

The complex factor  $A$  is

$$A = 3^{-1/3} \Gamma\left(\frac{2}{3}\right) (\sqrt{3}i - 1) = 1.63i - 0.94$$

For a perturbation  $\rho_1 \propto e^{-i\omega s/c + ikz}$ , the dispersion relation gives dependence  $\omega$  vs  $k$

$$1 = -\frac{ir_0 c^2 Z(k)}{\gamma} \int \frac{d\delta (d\rho_0/d\delta)}{\omega + ck\eta\delta}$$

This is a standard formula for a coasting beam instability.

If bends do not fill the whole ring, we introduce a weighting factor in  $Z$

$$Z \rightarrow Z \frac{R}{\langle R \rangle}$$

where  $\langle R \rangle = C/2\pi$ .

For a Gaussian distribution function,  $\rho_0 = n_b(2\pi)^{-1/2} \exp(-\delta^2/2\delta_0^2)$

$$\frac{(kR)^{2/3}}{\Lambda} = -\frac{A}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \frac{dp p e^{-p^2/2}}{\Omega + p}$$

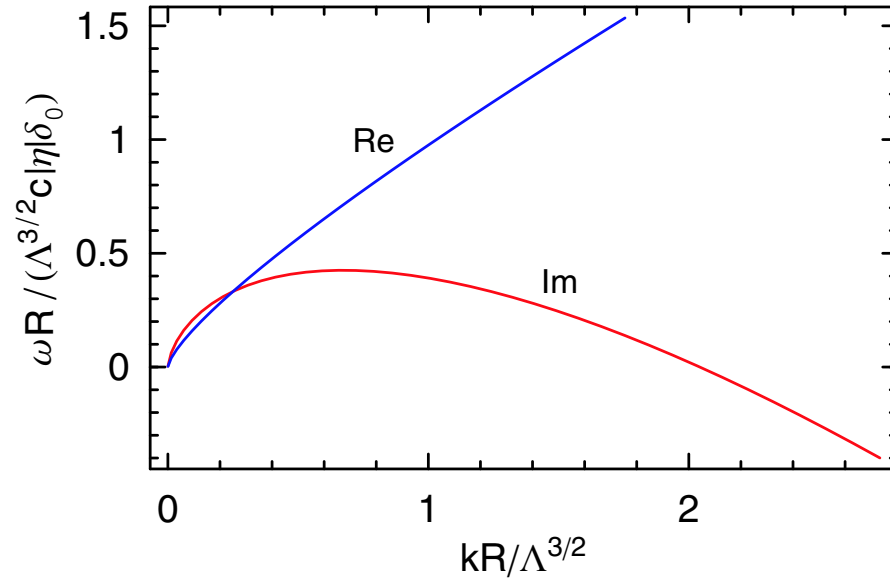
where  $\Omega = \omega/ck\eta\delta_0$ ,

$$\Lambda = \frac{n_b r_0}{|\eta| \gamma \delta_0^2} \frac{R}{\langle R \rangle}$$

Note

$$n_b r_0 = \frac{I}{17 \text{ kA}}$$

$\omega$  versus  $k$  for positive  $\eta$



The beam is unstable for such wavelength that

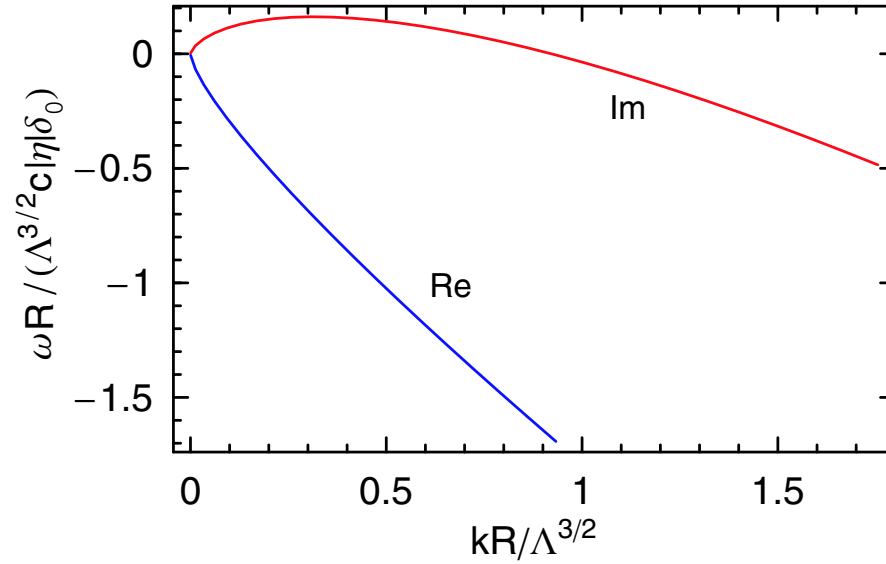
$$kR < 2.0\Lambda^{3/2}.$$

$$\Lambda \propto \frac{I}{\delta_0^2}$$

The maximum growth rate is reached at  $kR = 0.68\Lambda^{3/2}$  and is equal to  $\text{Im}\omega_{\text{max}} = 0.43\Lambda^{3/2}c\eta\delta_0/R$ .



$\omega$  versus  $k$  for negative  $\eta$



The beam is unstable for such wavelength that

$$kR < 0.92\Lambda^{3/2}.$$

The maximum growth rate is reached at  $kR = 0.32\Lambda^{3/2}$  and is equal to  $\text{Im}\omega_{\text{max}} = 0.16\Lambda^{3/2}c|\eta|\delta_0/R$ .

Limit of the cold beam,  $k \ll \Lambda^{3/2}/R$ ,

$$\text{Im } \omega = 1.2c \left( \frac{r_0 k^{4/3} n_b \eta}{\gamma R^{2/3}} \right)^{1/2}$$

No energy spread  $\delta_0$  in this formula.

We neglected the synchrotron damping  $\gamma_d$  due to incoherent radiation. The effective growth rate of the instability  $\sim \text{Im } \omega - \gamma_d$  (see details in Heifets and Stupakov, SLAC-PUB-8803).

**If we want to see the instability ...**

**1. Bunched Beam.** For a bunched beam of length  $\sigma_z$  the coasting-beam approximation can be applied if  $k\sigma_z \gg 1$ ,

$$\sigma_z \gtrsim 0.5 \frac{R}{\Lambda^{3/2}}$$

**2. Shielding.** Finite aperture  $a$  of the beam pipe — CSR is suppressed due to the shielding effect at

$$kR \lesssim \left( \frac{\pi R}{2a} \right)^{3/2}$$

Hence the instability can only develop for such values of  $k$  that  $2.0\Lambda^{3/2} > kR \gtrsim (\pi R/2a)^{3/2}$ .

$$\frac{R}{a} \lesssim \Lambda.$$

## Numerical Estimates for LER, ALS and VUV rings

Accelerator	LER PEP-II	ALS	VUV NSLS
$E$ (GeV)	3.1	1.5	0.81
$\eta$	$1.31 \cdot 10^{-3}$	$1.41 \cdot 10^{-3}$	$2.35 \cdot 10^{-2}$
$\delta_0$	$8.1 \cdot 10^{-4}$	$7.1 \cdot 10^{-4}$	$5.0 \cdot 10^{-4}$
$\langle R \rangle$ (m)	350	31.3	8.11
$R$ (m)	13.7	4	1.91
$a$ (cm)	2	1	2.1
$I_b$ (mA)	2	7.6 (30)	400
$\sigma_z$ (cm)	1	0.7	4.7
$\Lambda$	7	306 ( $1.2 \cdot 10^3$ )	250
$R/a$	550	400	90
$R/2\Lambda^{3/2}$ (cm)	1.0	$0.037(4.7 \cdot 10^{-5})$	0.025

## Discussion—formation length and retardation

We neglected retardation in the Vlasov equation. This is valid if the wake formation time is much smaller than the inverse growth rate of the instability:

$$t_{\text{form}} \sim \frac{R}{c} \frac{1}{(kR)^{1/3}} \ll \frac{1}{\text{Im } \omega}.$$

Using for the characteristic wavenumber and frequency of the instability  $kR \sim \Lambda^{3/2}$  and  $\text{Im } \omega \sim \Lambda^{3/2} c |\eta| \delta_0 / R$  yields the condition of applicability of the theory

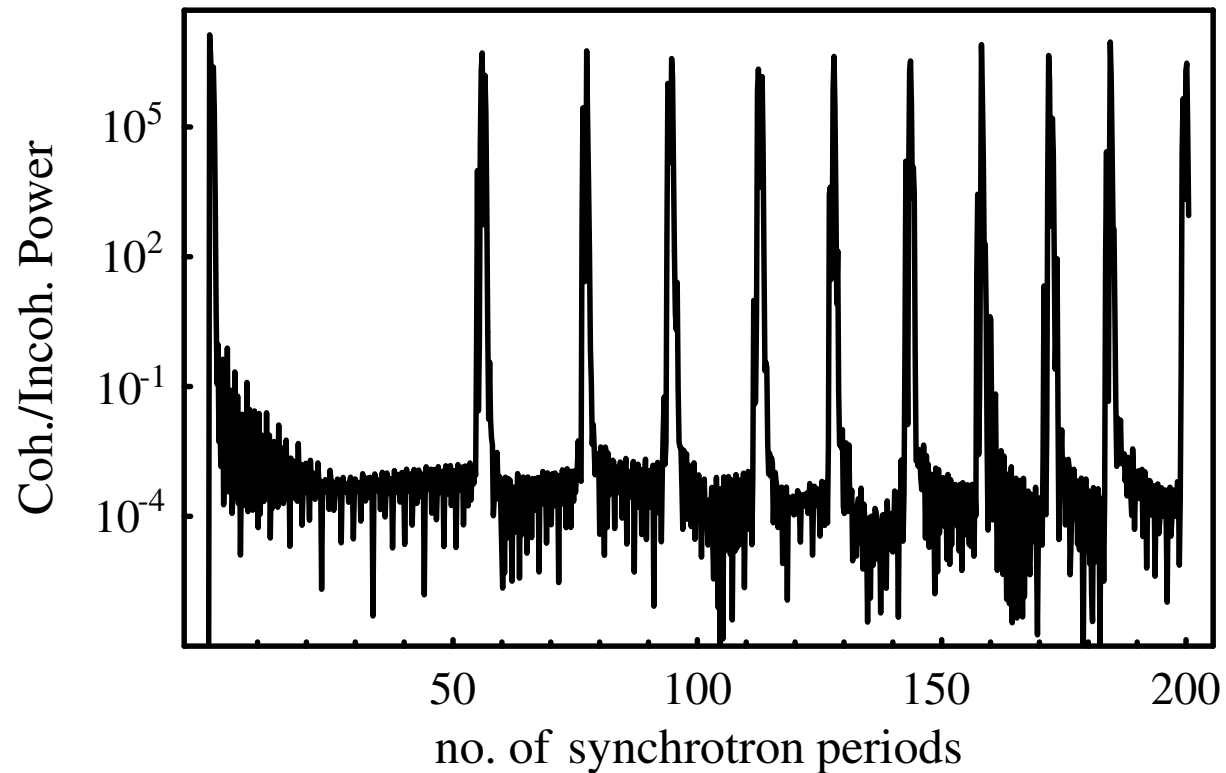
$$\frac{n_0 r_0}{\delta_0 \gamma} \ll 1.$$

### **Discussion—finite transverse size of the bunch**

The CSR wake is not applicable for very short wavelength. This wake was derived for a bunch that is infinitely thin in the transverse direction and assumes that all particles in the cross section of the bunch radiate coherently. However, the transverse coherence length  $l_{\perp} \sim k^{-2/3} R^{1/3}$  decreases with the wavelength and for very large values of  $k$  becomes smaller than the transverse dimension of the beam.

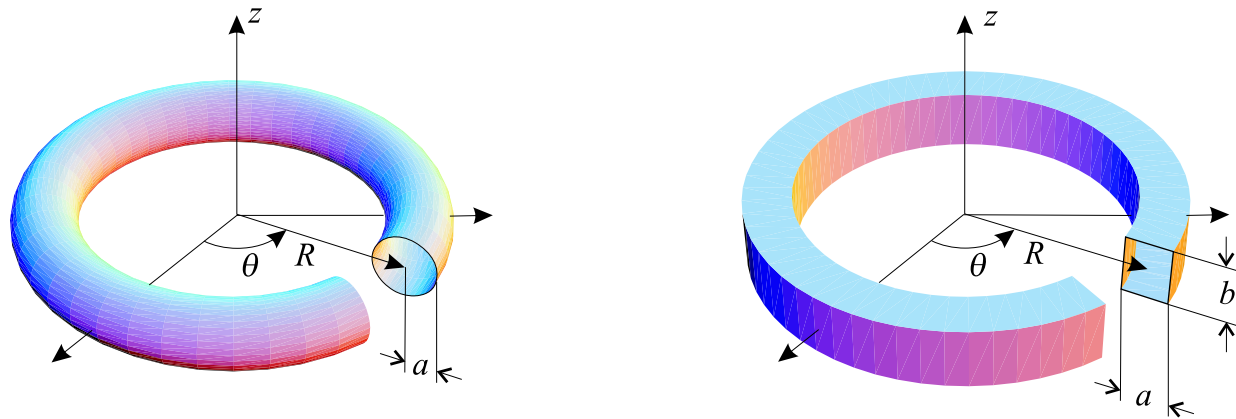
## Nonlinear Regime

Not much can be done analytically. See paper by M. Venturini and R. Warnock, SLAC-PUB-9505.



## Shielding and discrete modes

Near the shielding threshold,  $\lambda \sim a^{3/2}/R^{1/2}$ , the vacuum CSR impedance is not applicable. In the model of a toroidal waveguide with perfectly conduction walls and circular orbit, there are discrete synchronous modes that interact with the beam (B. Warnock&P. Morton, K.-Y. Ng, et al.). We did a new analysis of the shielded CSR impedance (G. Stupakov and I. Kotelnikov, SLAC-PUB-9553), which deals with arbitrary shape of the toroid cross section.

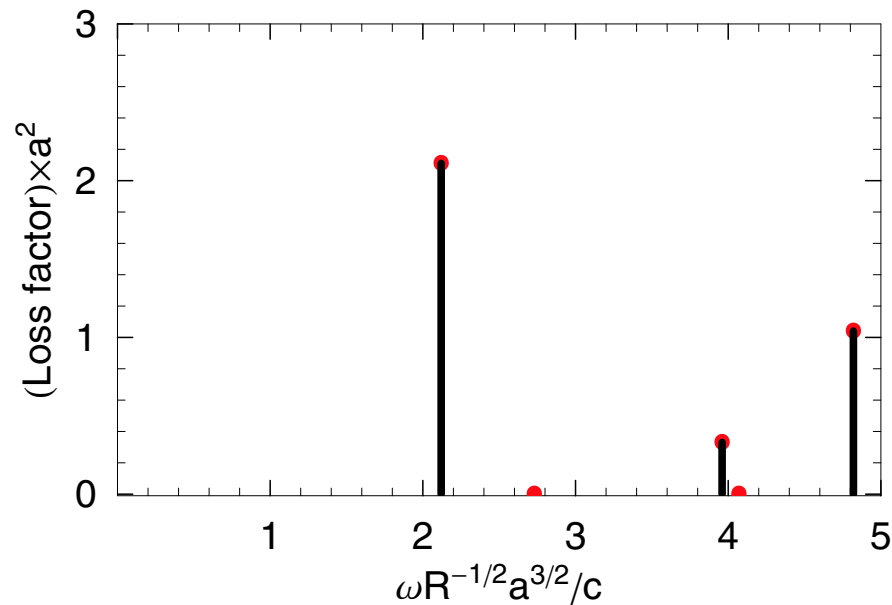




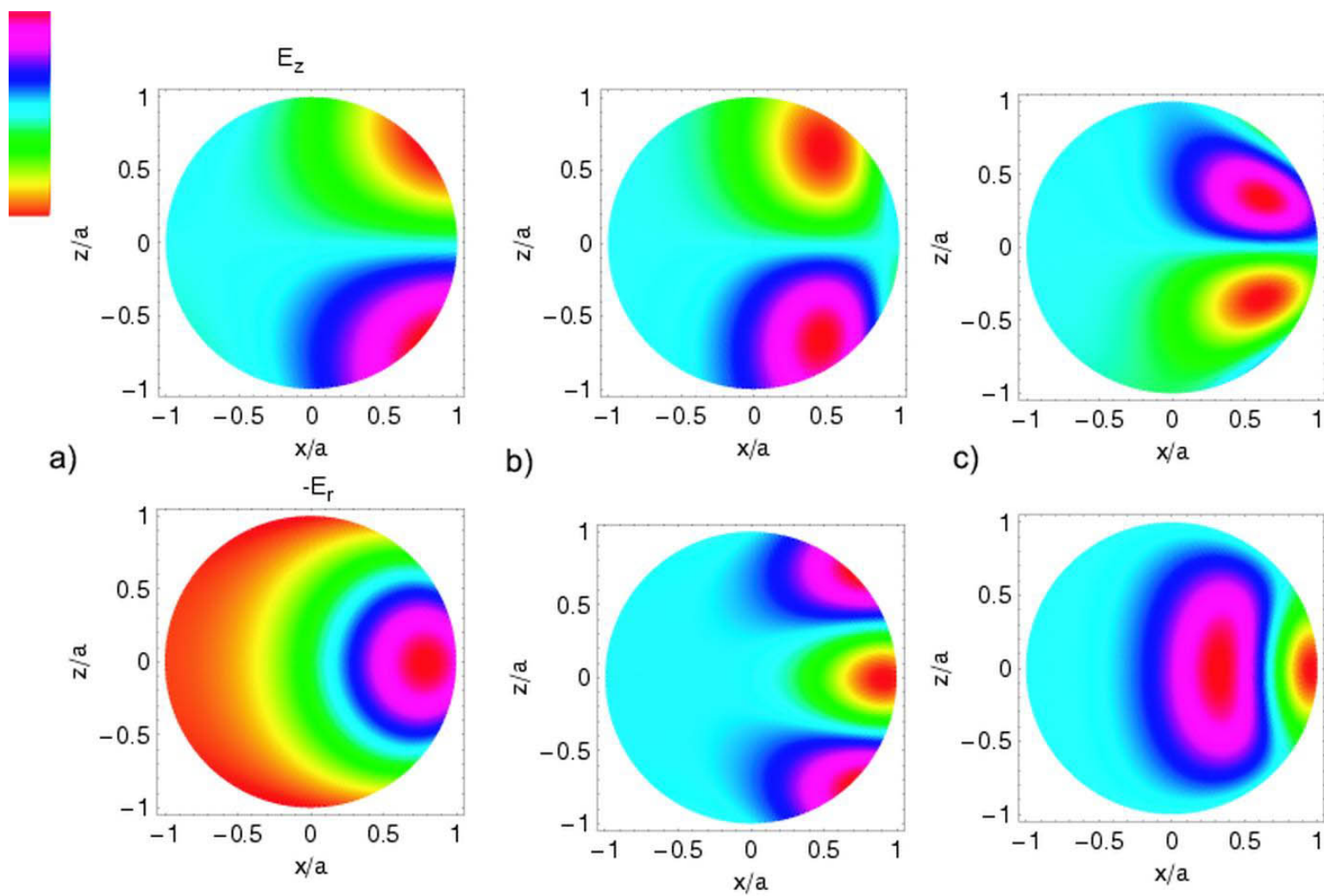
Each mode is characterized by the loss factor (per unit length)

$$w(z) = 2\kappa \cos\left(\frac{\omega}{c}z\right)$$

Fig. below—loss factors for a round toroid.

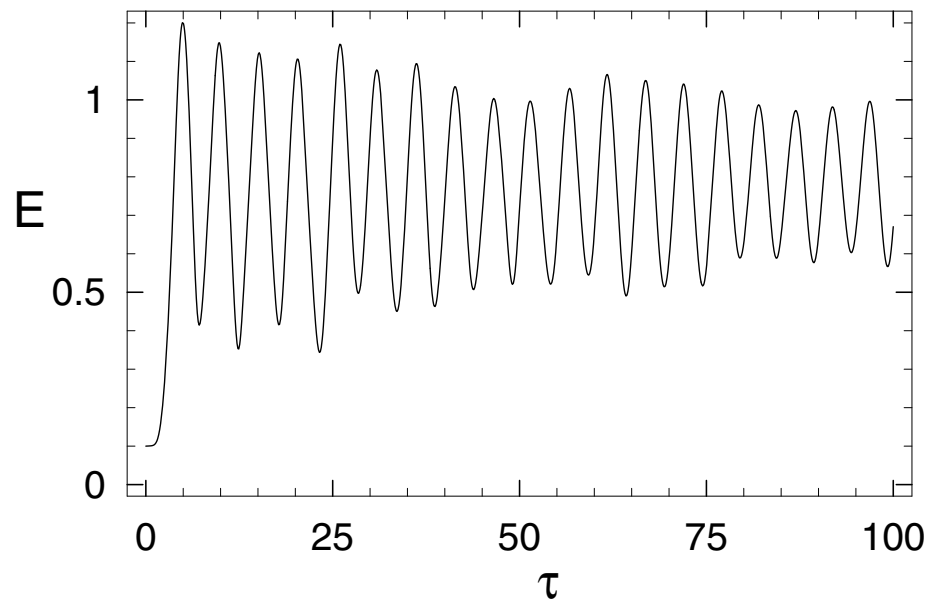


For  $R = 4$  m and  $a = 1$  cm, the unit of frequency ( $\omega/2\pi$ ) on the plot is 95 GHz, the unit of loss factor is 90 V/pC/m.



CSR instability should be treated as interaction with single modes, not a continuous spectrum (talk by S. Heifets). The theory is similar to 1D SASE FEL instability (the equations in scaled variables are identical)

$$\Gamma \sim I^{1/3}$$

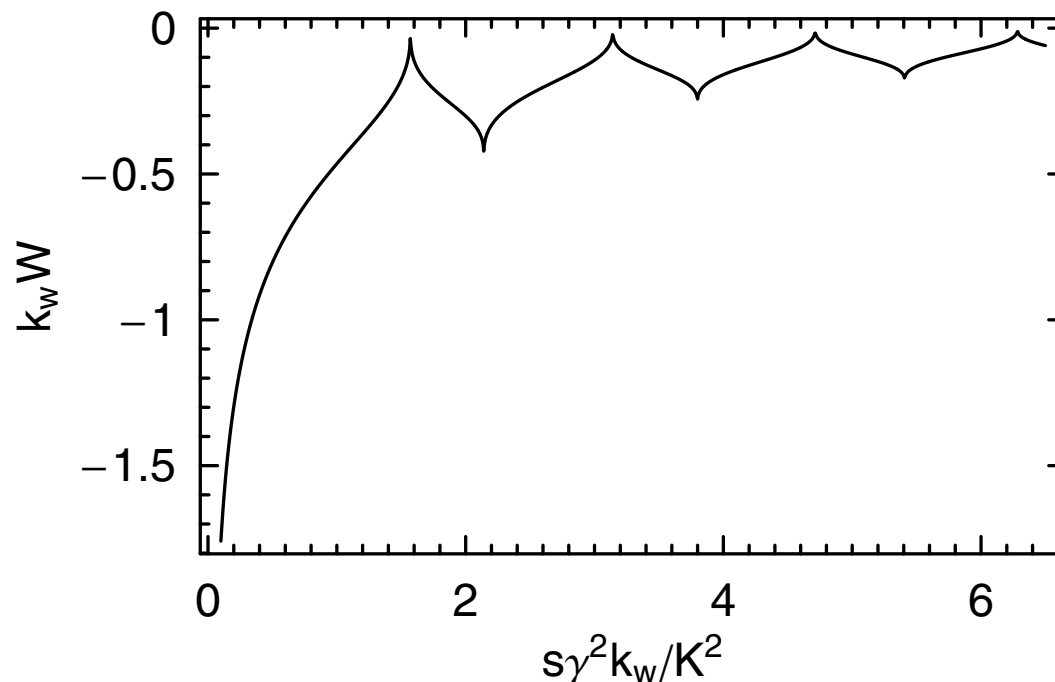


The promise of this theory is a possibility of quasi-continuous radiation (no bursts).

## CSR instability in wiggler

Sometime rings have wigglers. Example: NLC damping ring— $B = 2.15$  T,  $\lambda_w = 27$  cm,  $L_w = 46$  m,  $C = 300$  m. How does wiggler effect the CSR instability? We need the CSR wake for the undulator.

Wiggler wake potential in the limit  $K^2/2 \gg 1$ .



Talk by J. Wu.

## Conclusion

1. The theory of CSR instability in the ring is under rapid development. Areas of research include:
  - Single mode CSR instability—both linear and nonlinear regimes
  - Instability in a ring with wigglers
  - Shielding effect, arbitrary cross section of the vacuum chamber
  - Computer simulation—finite bunch length, shielding, nonlinear effects, radiation damping and diffusion
2. I see the challenge in searching for regimes where the beam is unstable, but the instability saturates into a steady state with a large CSR radiation.